# ANALYTICAL SOLUTION OF THE SCHRÖDINGER EQUATION WITH THE QUASI-HARMONIC POTENTIAL AND CENTRIFUGAL TYPE TERM VIA LAPLACE TRANSFORM

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Abstract. We consider the Schrödinger equation (*SE*) with the potential:  $V_q(r) = \delta r^2 + \frac{A}{r} + \frac{B}{r^2}$  defined by us as the quasi-harmonic potential with  $\frac{B}{r^2}$  the centrifugal type term,  $\delta, A, B > 0$  and  $0 < \mu \delta \ll 1$ . Applying Laplace transform method (*LTM*), we obtain a new analytical solution in the  $V_q$ - potential problem. Using directly and inverse Laplace transforms, we give the complete forms of the energy eigenvalues and the wave functions. Furthermore, introducing the potentials family  $\{\lambda V_q\}_{\lambda>0}$ , we outline a path for deriving the critical value of the angular momentum  $\ell_c$  depending on the scalar minimum value  $\lambda_c$  chosen such that bound states exist. For this family of potentials, we obtain a useful approximation of upper bound  $\ell_c^+$  to  $\ell_c$ .

*Key words:* Schrödinger equation, Laplace transform, analytical eigenfunctions, quasi-harmonic potential with centrifugal term

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# **1. INTRODUCTION**

In quantum mechanics, several authors have solved the Schrödinger equation [1] for potentials, such as pseudo harmonic, Mie type, Coulomb-like potentials [2–4], and Halton, Manning Rosen [5,6], Pöshel-Teller [7] and Wood-Saxon potential at the nuclear scale [8,9].

There are several methods for solving the *SE* such as Laplace transform [10], Nikiforov-Uvarov [6], homotopy perturbation [11], series solution [12], Fourier transform [13], asymptotic iteration methods [14], super-symmetric approach [15], variational method [16] and others.

In this paper, we note that the Laplace transform leads to the analytical and exact forms of eigenfunctions for the potential:

$$V_q(r) = \delta r^2 + \frac{A}{r} + \frac{B}{r^2},\tag{1}$$

which is a special quasi-harmonic type potential containing a centrifugal type term.

The  $V_q$  potential at the nuclear scale implies a short range behaviour of the potential. For this reason, there is a finite number of bound states beyond which the  $\ell$ -state is unbounded [17]. It is important to obtain the critical value of angular momentum  $\ell_c$ , defined indirectly by  $\lambda > 0$ , a scalar chosen to ensure  $V_{eff} < 0$  (see to Sect. 3.2), which gives rise to bound states for the short-range potential. Therefore, the second aim of this paper is to obtain an approximation for  $\ell_c$ . More exactly, in the case of  $V_q$  potential, we find an upper bound  $\ell_c^+$  to  $\ell_c$ .

The present paper gives an analytical solution to the  $V_q$ - potential problem by *LTM* to solve SE. In Section 2, we find the proper Laplace transform to solve SE. In Section 3, we give two categories of results: in subsection 3.1, the energy and eigenfunctions, and in subsection 3.2, the assessment of  $\ell_c^+$ ; we also introduce a family of  $V_q$  potentials, determined by some values of scalar  $\lambda$  to ensure the bound states.

Conclusions close our paper.

## 2. THE PROPER LAPLACE TRANSFORM FOR BOUND STATES SPECTRUM

In the natural units  $c = \hbar = 1$ , assuming spherical symmetry of the potential, the time-independent Schrödinger equation in the spherical coordinates  $(r, \theta, \phi)$  is given by [6, 18]:

$$\left[-\frac{1}{2\mu}\nabla^2 + V(r)\right]\Psi_{n\ell m}(r,\theta,\varphi) = E_{n\ell}\Psi_{n\ell m}(r,\theta,\varphi),$$
(2)

where  $E_{n\ell}$  and V(r) denote the energy eigenvalues, the potential and  $\mu$  the reduced mass, respectively, for a certain physical system.

The  $\Psi_{n\ell m}(r, \theta, \phi)$  denotes the n-th state of eigenfunctions. We choose the bound state eigenfunctions  $\Psi_{n\ell m}(r, \theta, \phi)$  such that wave functions are vanishing for  $r \longrightarrow 0$ ,  $r \longrightarrow \infty$ . We look for separable solutions form of *SE*:

$$\Psi_{n\ell m}(r,\theta,\phi) = \Re_{n\ell}(r)Y_{\ell}^{m}(\theta,\phi), \qquad (3)$$

where  $\Re_{n\ell}(r)$  are the radial functions and  $Y_{\ell}^{m}(\theta, \varphi)$  the angular functions, respectively.

Equation (2) provides two separated equations: one is known as the spherical harmonics equation and the other is known as the radial equation, which generally takes the following form in a *D*- dimensional space with the hyper-sphere  $\Sigma_D = \Sigma_D(r, \varphi, \theta_1, ..., \theta_{D-2})$ :

$$\Re''(r) + \frac{2}{r}\Re'(r) + \left[-\frac{\ell(\ell+D-2)}{r^2} + 2\mu(E_{n\ell} - V(r))\right]\Re(r) = 0,$$
(4)

where  $\ell(\ell + D - 2)$  is the separation constant with D > 1 and  $\ell = 0, 1, 2, 3, ..., n - 1$ .

In this paper, because we consider the stationary case of SE with the potential  $V_q$ , we work in three spatial dimensions D = 3. So, eq. (4) becomes:

$$\Re''(r) + \frac{2}{r}\Re'(r) + \left[-\frac{\ell(\ell+1)}{r^2} + 2\mu\left(E_{n\ell} - \delta r^2 - \frac{A}{r} - \frac{B}{r^2}\right)\right]\Re(r) = 0.$$
(5)

In the eq. (5), for  $r \to \infty$ , we consider the following asymptotic form: [19]

$$\mathfrak{R}''(r) - 4d^2r^2\mathfrak{R}(r) = 0,$$

$$d = \sqrt{\frac{\mu\delta}{2}}, \ \delta > 0,$$
(6)

and consequently, we propose to find the solution in the following form:

$$\Re(r) = r^k e^{-dr^2} f(r), \, k > 0.$$
(7)

Further, we solve the following SE form with k-value and f(r) function as unknowns:

$$r^{2}f''(r) + r(\eta_{k} + 2r - 2dr^{2})f'(r) + \left[Q_{n\ell} - 2\mu Ar + \epsilon_{k}r^{2} + dkr^{3} - 2\mu\delta r^{4} + 4d^{2}r^{4}\right]f(r) = 0,$$
(8)

where the prime over f(r) denotes the derivative with respect to r; also, we introduce the following notations:

$$Q_{n\ell} = k(k+3) - \ell(\ell+1) - 2\mu B, \in_k = 2\mu E_{n\ell} - 4dk - 6d, \eta_k = 2k.$$
(9)

Starting from this point, we impose several parametric restrictions. Firstly, we aim to obtain:

$$Q_{n\ell} = 0. \tag{10}$$

For that, we consider the following correlation  $\mu B = \ell$ , which is possible physical speaking, because both  $\mu > 0$  (as a mass) and B > 0, with  $\mu B$  taking an integer value like  $\ell$ ; also, it is useful from a mathematical point of view, because it simplifies the  $Q_{n\ell}$  expression; after that, the eq.(10) implies two values for k:

$$k_{+} = \ell,$$
  
 $k_{-} = -(\ell + 3).$ 
(11)

The acceptable physical value remains  $k_+ = \ell$ . For r > 0, the equation (8) becomes:

$$rf''(r) + (\eta_{\ell} + 2r - 2dr^{2})f'(r) + [4d^{2}r^{3} - 2\mu\delta r^{3} + d\ell r^{2} + \epsilon_{\ell}r - 2\mu A]f(r) = 0.$$
(12)

In eq. (12), in order to reduce the complexity of its polynomis coefficients, we simultaneously consider two restrictions: a) a parametric restriction  $\mu\delta \ll 1$  and b) a scale restriction  $r \ll 1$ , making d and r small enough (e.g.  $1 \sim a_0$  - the first Bohr radius at the atomic scale, or  $1 \sim$  the nucleus radius at the nuclear scale, depending on the considered physical conditions).

Hence, the following differential equation in the unknown function f(r) remains unsolved:

$$rf''(r) + (\eta_{\ell} + 2r - 2dr^2)f'(r) + (d\ell r^2 + \epsilon_{\ell} r - 2\mu A)f(r) = 0.$$
(13)

So, in eq. (8), before working in the transform space by applying Laplace transform, we reduce the differential equation from third to the second order by using  $Q_{n\ell} = 0$  with the above restrictions.

Therefore, we apply the Laplace transform  $\Phi(s) = L\{f(r)\}(s)$ , with Re(s) > 0 [10], and eq. (13) becomes:

$$d(2s-\ell)\Phi''(s) + (s^2+2s+\tilde{\epsilon})\Phi'(s) + (\gamma s+\alpha)\Phi(s) = 0, \tag{14}$$

where :

$$\begin{aligned} \gamma &= 2 - 2\ell, \\ \alpha &= 2\mu A + 2, \\ \tilde{\epsilon} &= \epsilon + 4d. \end{aligned}$$
 (15)

We observe that  $s_0 = \frac{\ell}{2}$  is a singular point, suggesting the following form of the Laplace transform:

$$\boldsymbol{\Phi}(s) = \frac{C_{n\ell}}{\left(s - \frac{\ell}{2}\right)^{n+1}}.$$
(16)

The inverse Laplace transform  $f(r) = L^{-1} \{\Phi(s)\}(r)$  leads to the following expression:

$$f(r) = \frac{C_{n\ell}}{n!} r^n e^{\frac{\ell r}{2}}.$$
 (17)

#### **3. RESULTS**

## 3.1. The energy eigenvalues and wave functions

Using the form (16) in eq.(14), we obtain the system of conditions:

$$\begin{split} \gamma &= n+1, \\ \alpha &= 2(n+1) + d\ell\gamma, \\ \tilde{N} - \ell\alpha &= 2(n+1)\tilde{\in}, \end{split}$$
 (18)

where  $\tilde{N} = 4d(n+1)(n+2)$ .

Solving the third equation from (18), we obtain the energy eigenvalues:

$$E_{n\ell} = \frac{d}{\mu} \left[ n + 2\ell + 3 - \frac{\alpha\ell}{4d(n+1)} \right] = \sqrt{\frac{\delta}{2\mu}} \left[ n + 2\ell + 3 - \ell \frac{(4+\ell)}{8d} \right].$$
 (19)

We remark that in the case of harmonic oscillator  $V(r) = \frac{1}{2}\mu\omega^2 r^2$ , the energy expression leads to the eigenvalue  $E_{00} = \frac{3}{2}\omega$  which is the well known ground state energy.

A complete solution of SE implies the computation of normalization constant  $C_{n\ell}$  from the condition:

$$\int_0^\infty \left[\Re(r)\right]^2 r^2 dr = 1.$$
 (20)

In order to calculate elegantly the normalization constant, we propose to use a special and remarkable integral; but for that, it is necessary to take in account the approximation  $r - \frac{\ell}{4d} \approx r$  which occurs in the exponential part of the integral by the above condition (20).

The remarkable integral is:

$$\int_{0}^{\infty} x^{p} e^{-ax^{q}} dx = \frac{1}{q} \frac{\Gamma\left(\frac{p+1}{q}\right)}{a^{\frac{p+1}{q}}}, p, q, a > 0,$$
(21)

where  $\Gamma(\cdot)$  is the gamma function.

Thus, we obtain the normalization constant:

$$C_{n\ell} = n! \left\{ \frac{2(2d)^{\ell+n+\frac{3}{2}}}{\Gamma(\ell+n+\frac{3}{2})} e^{-\frac{\ell^2}{8d}} \right\}^{\frac{1}{2}}.$$
(22)

Finally, using the relations (11), (17), (22) in the radial function (7), we find the complete analytical form of the eigenfunctions:

$$\Re_{n\ell}(r) = r^{n+\ell} \frac{C_{n\ell}}{n!} e^{-dr^2 + \frac{\ell}{2}r}.$$
(23)

Further, we compute the  $r_{rms}$  radius:

$$r_{rms} = \sqrt{\langle r^2 \rangle},\tag{24}$$

where

$$< r^{2} >= \int_{0}^{\infty} r^{2} \left[ \Re(r) \right]^{2} r^{2} \mathrm{d}r,$$
 (25)

and obtain :

$$r_{rms} = \sqrt{\frac{\ell + n + \frac{3}{2}}{2d}} = \sqrt{\frac{\ell + n + \frac{3}{2}}{\sqrt{2\mu\delta}}}, \delta \neq 0.$$
(26)

## 3.2. The upper bound approximation of angular momentum

The methodology for the *SE* via *LTM* is also useful in the short range potential case [20]. At the nuclear scale, we introduce the family of potentials  $\{\lambda V_q\}_{\lambda>0}$  with  $\delta, A, B > 0$ .

In eq. (4), considering the transformation  $\Re(r) = \frac{U(r)}{r}$  and unit  $2\mu = 1$ , we obtain a SE form [21]:

$$\left[-\frac{d^2}{dr^2} - \lambda V_q(r) + \frac{\ell(\ell+1)}{r^2}\right] U_{n\ell}(r) = \tilde{E}_{n\ell} U_{n\ell}(r).$$
(27)

Combining  $\lambda > 0$  and  $V_q$  at nuclear scale, the  $-\lambda V_q$  term is an attractive potential with  $\lambda$  as a measure of the  $V_q$ - family's strength.

We consider the effective potential: [17]

$$V_{eff}(r) = -\lambda V_q(r) + \frac{\ell(\ell+1)}{r^2}.$$
 (28)

A possible bound state of positive energy corresponds to the proposed  $\lambda V_q$  potentials including quasibound states, where  $\lambda$  is relevant in the binding of  $\ell$ -states.

An infinitely small but negative part of  $V_{eff}$  would permit a bound state.

So, the critical strength value  $\lambda_c(\ell)$  is needed to bind a  $\ell$ -state, thus occurring indirectly the critical value for  $\ell$ .

Regarding the above, we make several computations:

• We define  $r_0$ - the radius such that  $V'_{eff}(r_0) = 0$  and we obtain:

$$r_0 = \sqrt[3]{\frac{A}{2\delta}} > 0, \delta \neq 0.$$
<sup>(29)</sup>

• The critical strength value of  $\lambda_c$  occurs as the minimum value necessary to get a bound state from the condition:

$$\lambda_{c} \ge \ell(\ell+1) \frac{2}{-r_{0}^{3} V_{q}'(r_{0})}, V_{q}'(r_{0}) < 0.$$
(30)

• We obtain the upper bound  $\ell_c^+$  as:

$$\ell_c^+ \approx \sqrt{\frac{-r_0^3 V_q'(r_0)}{2}} \sqrt{\lambda_c} = \sqrt{\frac{1}{2} (2B + (4^{-\frac{2}{3}} - 1)Ar_0)} \sqrt{\lambda_c},$$
(31)

this upper bound approximation being appropriate for practical work.

# 4. CONCLUSIONS

In the two body problem, associated with a  $V_q$  quasi-harmonic potential having  $0 < \mu \delta \ll 1$ , and our defined parametric constraints, we solved 3-dimensional Schrödinger equation via Laplace transform method. We obtained a complete analytical solution, namely the energy eigenvalues and eigenfunctions.

In the computation of k positive value, we considered the correlation  $\mu B = \ell$  and obtained  $k_+ = \ell$ . Therefore, the k value, involved in parametric restriction  $Q_{n\ell} = 0$ , may be a subject of interest in quantum computing.

Using the wave function, we computed the root-mean-square (*rms*) charge radius, which is measured for most stable nuclei by electron scattering form factors and/or from the *x*-ray transition energies of muonic atoms. It will be interesting to compare our formula of  $r_{rms}$  with experimentally-derived values.

Considering at nuclear scale a potentials family  $\{\lambda V_q\}_{\lambda>0}$ , such that  $-\lambda V_q$  term becomes an attractive potential, we also obtained the analytical solution, which is only valid for well bound states, but not for angular

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momentum close to or above the critical  $\ell_c$ . We estimated the  $\lambda_c(\ell)$ , representing the necessary strength of the  $\lambda V_q$  potentials, required such that  $V_{eff}$  contains a negative part. Therefore, we calculated an upper bound  $\ell_c^+$  which is a good approximation of  $\ell_c$  and find  $\ell_c^+$  is proportional to  $\sqrt{\lambda_c}$ , namely with the strength of the potentials family  $\{\lambda V_q\}_{\lambda>0}$ .

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